

Accuracy of ETC Calculations

This version of the ETC is intended to help users plan observations for General Observer (GO) programs. The ETC approximates our current best knowledge and understanding of the performance of the JWST instruments, based on inflight measurements and calibrations. It has been validated against inflight observations analyzed by the instrument teams. Users should exercise appropriate caution when interpreting results from the ETC. A number of known issues remain, which may affect predicted sensitivities.

See Known Issues for additional details.

The ETC is not intended to be a complete observation simulator, and some higher-order effects are not accounted for, such as field distortion.



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01 - What is the ETC?

Exposure Time Calculator

- Main end-product is the **detector configuration** you will use

Choosing the instrument:

Depends on your science case...

And the instrument capability (sensitivity, etc)

For example:

- You need 7 18 microns spectroscopy
- go to : https://jwst-docs.stsci.edu/ (or remember last presentation)
 - MIRI time!

So either Low-Resolution or Mid-Resolution

- The **higher** the resolution the **more** photons needed

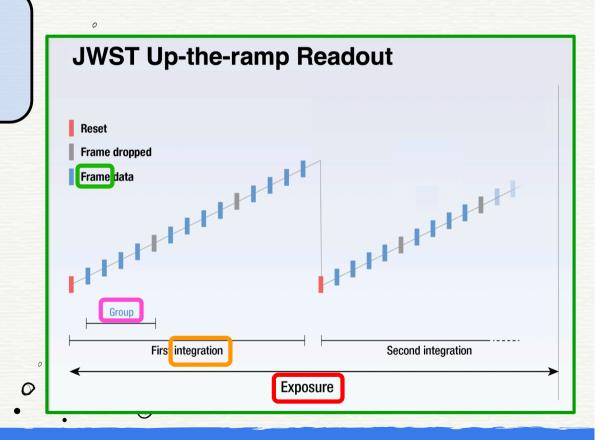
01 - What is the ETC?

Exposure Time Calculator

- Main end-product is the **detector configuration** you will use

Detector configuration i.e. the number of :

- Frame per group
- Groups per integration
- Integration per exposure
- Exposures



01 - What is the ETC?

Under the hood: The ETC engine

The JWST ETC engine uses a <u>pixel-based 3-dimensional</u> approach to perform calculations on small (typically a few arcseconds) 2-dimensional <u>user-created astronomical scenes</u>. It models both the <u>spatial and the wavelength dimensions</u>, using realistic point spread functions (produced using <u>WebbPSF</u>) for each instrument mode. It natively handles <u>correlated read noise</u>, inter-pixel capacitance, and <u>saturation</u>. Since the signal and noise are modeled for individual detector pixels, the ETC is able to replicate many of the steps that observers will perform when calibrating and reducing their JWST data. This simplifies interpretation of the extracted signal-to-noise ratio (SNR) calculated by the ETC.



02 - Usefulness?

Quick and dirty way to check feasibility & estimate time

Thorough proposal preparation

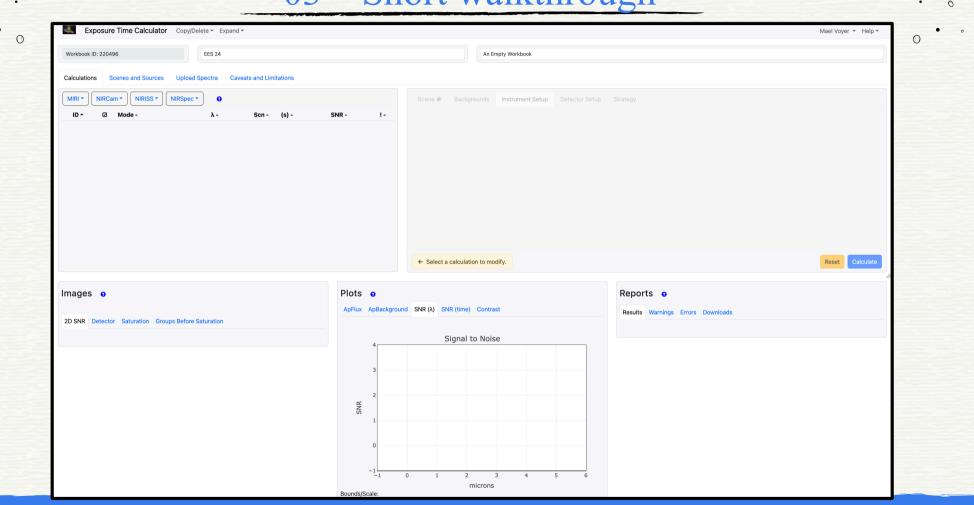
10 min time:

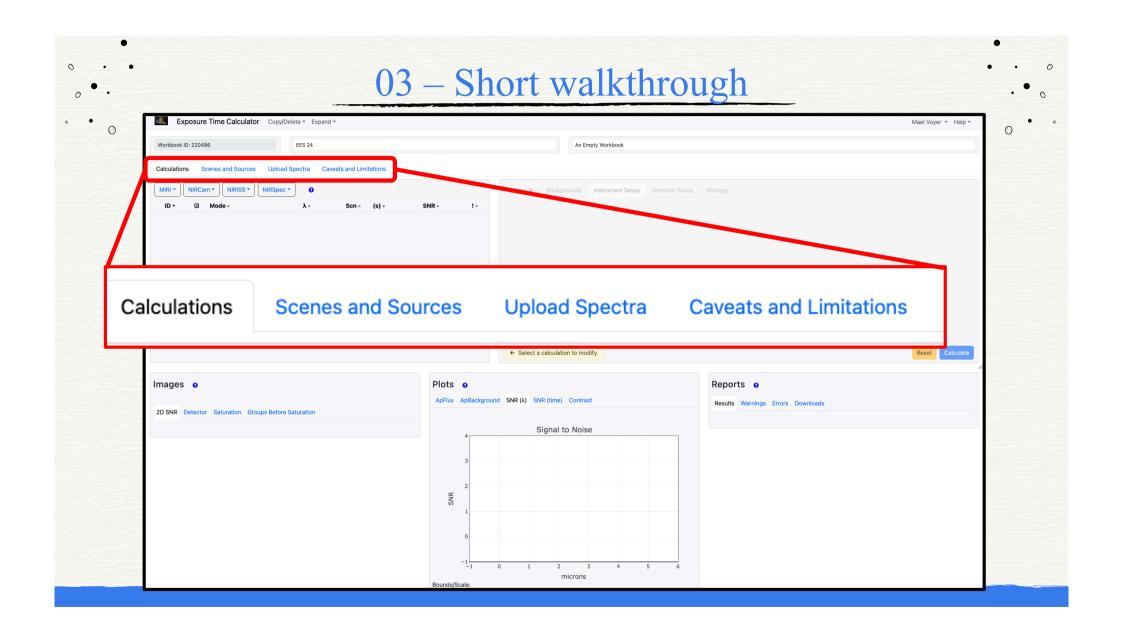
- Idea of the spectra
- Some flux estimate

Much more time :

- Detailed parameters
- Check and re-check
- Complementary to the APT

03 – Short walkthrough



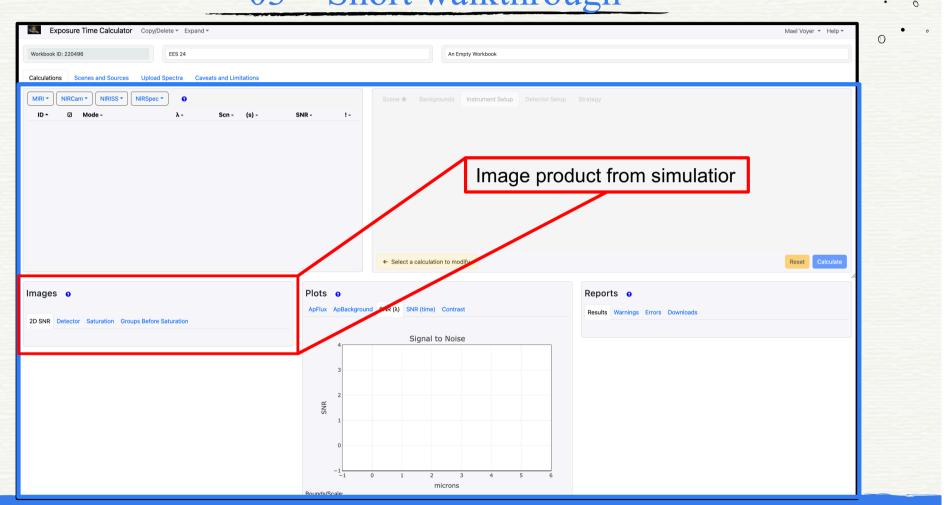


03 – Short walkthrough Exposure Time Calculator Copy/Delete * Expand * Mael Voyer ▼ Help ▼ Instrument modes for calculation Reset Calculate Plots 9 Reports 9 Images 9 ApFlux ApBackground SNR (λ) SNR (time) Contrast Results Warnings Errors Downloads 2D SNR Detector Saturation Groups Before Saturation Signal to Noise

03 – Short walkthrough



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03 – Short walkthrough Exposure Time Calculator Copy/Delete - Expand -Mael Voyer - Help -Scn - (s) -Analysis of the simulator results Reset Calculate ← Select a calculation to modify. Images 0 Plots 9 Reports 9 ApFlux ApBackground SNR (λ) SNR (time) Contrast Results Warnings Errors Downloads 2D SNR Detector Saturation Groups Before Saturation Signal to Noise

03 – Short walkthrough Exposure Time Calculator Copy/Delete - Expand -Mael Voyer - Help -Scn - (s) -Analysis of the simulation & errors Reset ← Select a calculation to modify. Images 0 Plots 9 Reports 9 ApFlux ApBackground SNR (λ) SNR (time) Contrast Results Warnings Errors Downloads 2D SNR Detector Saturation Groups Before Saturation Signal to Noise

Thank you

